

↑ [HF radio propagation](#):

Scientific and Technical Concepts You Should Know

This page is part of the project "[Understanding HF Propagation](#)."
by [Doron Tal](#), [4X4XM](#)

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HF (High Frequency) radio propagation, used for long-distance communication, involves several key concepts and terminology. Knowing this terminology can help you understand long-distance communication via [skywaves](#).

Find below a glossary of [basic terms](#) and an [extended index of terms](#), all related to HF radio propagation:

A glossary of basic terms

[Waves](#) and [radio propagation](#)

1. [Frequency \(f\)](#) is the number of cycles of a radio wave that occur in one second, measured in hertz (Hz).
2. [Wavelength \(\$\lambda\$ \)](#) is the distance between two consecutive points in a wave, usually measured in meters (m).
The relationship between wavelength and frequency is inverse, with higher frequencies corresponding to shorter wavelengths and vice versa.
3. [Groundwave propagation](#) is the use of radio waves that travel along the earth's surface, providing communication over relatively short distances, typically up to a few hundred kilometers.
4. [Skywave propagation](#) is the phenomenon in which HF radio waves are bounced back to earth's surface, via the

ionosphere, allowing long-distance communication beyond the line of sight.

5. The [ionosphere](#) is the region of the earth's upper atmosphere (approximately 50 to 800 kilometers above the ground) where ionization occurs due to solar radiation. This region plays a crucial role in HF radio propagation by bouncing radio waves back to Earth. The ionosphere is continuously changing due to natural storms, as explained below:

A [storm](#) represents a disturbed state of the natural environment, characterized by significant disruptions to normal conditions. In this project, we specifically refer to:

- i. [Solar storms](#) occur on the sun.
- ii. [Space weather storms](#) occur in the space between the sun and Earth.
- iii. [Geomagnetic storms](#) (also known as magnetic storms) are temporary disturbances of the Earth's magnetosphere caused by a [solar wind](#) shock wave.
- iv. [Ionospheric storms](#) involve varying densities of energized free electrons in the ionosphere due to geomagnetic storms.

Understanding the dynamics of solar and space weather storms reveals the intricate relationship between our sun and Earth. Solar storms, originating on the sun, unleash powerful energy that travels through space, affecting not only the celestial environment but also our planet's magnetic field. When these solar winds encounter Earth's atmosphere, they can trigger geomagnetic storms, which are temporary disturbances that disrupt communication systems and power grids. Additionally, the presence of energized free electrons in the ionosphere gives rise to ionospheric storms, further complicating our technological landscape.

6. [Skip distance](#) is the minimum distance from the transmitter for receiving [skywaves](#), while [blind, or dead zone](#), refers to the region with poor or nonexistent reception.
7. [MUF](#) ([Maximum Usable Frequency](#)) is the highest frequency at which [skywave propagation](#) is reliably

achievable between two points on earth during a specific time and under specific [ionospheric conditions](#).

8. [LUF \(Lowest Usable Frequency\)](#) is the lowest frequency at which [skywave propagation](#) can be used effectively.
9. [Fading](#) (QSB) is the variation in signal strength and quality experienced during radio transmission due to changes in [ionospheric conditions](#), interference, or other factors.
10. The [F-region](#) is the upper region of the ionosphere that consists of two sub-regions, F1 and F2, and is the primary region for HF propagation. The F2 is particularly important for long-distance communication.
11. The [E-region](#) is the middle region located below the F-region. It is responsible for sporadic E propagation, which can enable HF communication at higher frequencies.
12. The [D-region](#) is the lowest region of the ionosphere, located at an altitude of approximately 48 to 90 kilometers. The D-region absorbs HF radio waves during daylight hours, affecting propagation.
13. The ionosphere bounces a radio wave at vertical incidence below the [critical frequency \(fc\)](#), regardless of transmitter power.
14. The [Optimum Working Frequency \(OWF\)](#) is the frequency that provides the best propagation conditions for a given HF communication path.
15. The [Solar Flux Index \(SFI\)](#) is a measure of the sun's radio emissions at a specific wavelength (usually 10.7 cm). It is used as an indicator of [solar activity](#), which affects [ionospheric conditions](#) and HF propagation.
16. The [A-index](#) is a daily measure of geomagnetic activity on a scale from 0 to 400, with lower values indicating quieter

geomagnetic conditions. Geomagnetic activity can disrupt HF propagation.

17. The [K-index](#) represents short-term (3-hour) geomagnetic activity at a specific geomagnetic station. It quantifies disturbances in Earth's horizontal magnetic field by comparing geomagnetic fluctuations, measured by a magnetometer, to a quiet day. The K-scale is logarithmic, a scale from 0 to 9.

Extended index of terms

A list of terms that are explained on this website:

A

1. [A-Index](#)
2. [Absorption of skywaves](#)
3. [ACE - Advanced composition explorer](#)
4. [ALE](#) — Automatic link establishment
5. [Amateur radio](#)
6. [Amateur radio bands](#)
7. [Amateur radio propagation today](#)
8. [Angle elevation/transmission](#)
9. [Applications for analyzing and forecasting HF propagation](#)
10. [Applications used for forecasts and predictions](#)
(references)
11. [Aurora](#)
12. [At what Kp index can I see aurora?](#)

B

13. [Band activity](#)
14. [Band conditions](#)
15. [Band conditions banners](#)

16. [Band Openings](#)
 17. [Beacons](#)
 18. [Beyond line of site](#) (BLOS)
 19. [Blackouts; R₁₋₅ scale](#)
 20. [Blind zone](#)
 21. [BLOS](#) (Beyond line of site)
 22. [Bt - total strength of the Interplanetary Magnetic field \(IMF\)](#)
 23. [Bz - orientation of the IMF](#)
-

C

24. [Carrington event](#)
25. [ChatGPT learns about forecasting tools for radio amateurs](#)
26. [Chirping](#)
27. [Clouds of plasma in the ionosphere](#)
28. [CME](#) - coronal Mass ejection
29. [Compare DR2W propagation tool with VOACAP](#)
30. [Compare forecasting tools for HF radio propagation](#)
31. [Compare forecasting tools for radio amateurs](#)
32. [Communication conditions](#)
33. [Complex HF propagation Modes](#)
34. [Conductive ground](#)
35. [Coronal Mass ejection \(CME\)](#)
36. [Current ham band conditions](#)
37. [Current ham band propagation](#)
38. [Current ham radio conditions](#)
39. [Current HF band conditions maps for radio operators](#)
40. [Current HF band conditions online map](#)
41. [Current HF bands conditions—charts for radio hams](#)
42. [Current HF propagation](#)
43. [Current HF propagation conditions](#)

44. [Current HF radio propagation](#)
 45. [Current propagation](#)
 46. [Current propagation conditions](#)
 47. [Current shortwave propagation conditions](#)
 48. [Current solar activity](#)
 49. [Current solar events](#)
 50. [Critical frequency](#)
 51. [Critical frequencies concept](#)
 52. [Critical frequency map](#)
-

D

53. [D-region](#)
 54. [Dead zone](#)
 55. [Digisonde directogram](#)
 56. [Direct wave propagation](#) (LOS - Line Of Site)
 57. [DRAP](#)
 58. [DX clusters](#) [DX Spots](#)
 59. [DX propagation map](#)
 60. [DX propagation online map](#)
 61. [DX propagation prediction](#)
 62. [DXlook](#)
 63. [DXMAPS](#)
 64. [DXWatch](#)
 65. [DXZone](#)
-

E

66. [E-region](#)
67. [Earth's magnetic field](#)
68. [Earth's Magnetosphere](#)
69. [Earth Magnetosphere vs Earth's magnetic field](#)
70. [Earth observing system \(EOS\)](#)
71. [Earth's magnetic field](#)

72. [Effects of solar activity on amateur radio](#)
 73. [Electron densities](#)
 74. [Electron flux \(Ef\)](#)
 75. [Electromagnetic Radiation](#)
 76. [Electromagnetic Spectrum](#)
 77. [Electromagnetic Wave](#)
 78. [Electromagnetic Spectrum](#)
 79. [Electromagnetic Radiation](#)
 80. [Electromagnetic Wave](#)
 81. [Elevation Angle](#)
 82. [EME Deg - earth-Moon-Earth Degradation](#)
 83. [EsEU - Sporadic e europe](#)
 84. [EsNA - Sporadic e North America](#)
 85. [EUV sunlight](#)
-

F

86. [FAQ: HF propagation](#)
87. [FAQ: Current HF propagation conditions](#)
88. [F10.7](#)
89. [F-region](#)
90. [Flare](#)
91. [f_oF₂ - critical frequency](#)
92. [foF2 map](#)
93. [Forecast HF propagation conditions](#)
94. [Forecast HF radio propagation - applications](#)
95. [Forecast propagation](#)
96. [Forecast propagation for radio amateurs](#)
97. [Forecast vs prediction of HF band conditions](#)
98. [Forecasting HF propagation for radio amateurs](#)
99. [Forecasting MUF for a 3000 km path](#)
100. [Forecasting tools for radio amateurs](#)
101. [Free Electron Density](#)

G

102. [G1-G5](#) - scale of Geomagnetic storms
 103. [Gamma rays](#)
 104. [Gama rays bursts](#) (GRB)
 105. [GRB](#) - Gama-rays bursts
 106. ["GeoMag"_\(index\) - earth`s agnetic field activity](#) GMF activity
 107. [Geomagnetic](#)
 108. [Geomagnetic activity](#) | [What is Geomagnetic activity](#)
 109. [Geomagnetic conditions](#)
 110. [Geomagnetic data](#)
 111. [Geomagnetic field activity](#)
 112. [Geomagnetic indices K, A](#)
 113. [Geomagnetic storms and HF communications](#)
 114. [Geomagnetic storms; G₁₋₅ scale](#)
 115. [Global conditions](#)
 116. [Global HF propagation](#)
 117. [Global radio propagation conditions](#)
 118. [Greyline](#)
 119. [Greyline radio propagation explained](#)
 120. [Greyline vs "Solar Terminator"](#)
 121. [Ground wave](#)
 122. [Ground wave propagation](#)
 123. [Ground waves](#)
-

H

124. [HAARP Project for Radio Hams](#)
125. [Hale cycle](#)
126. [Ham Activity](#)
127. [Ham band conditions](#)
128. [Ham bands' activity](#)
129. [Ham conditions](#)

130. [Ham propagation](#)
131. [Ham propagation map](#)
132. [Ham radio communications](#)
133. [Ham radio HF propagation \(real-time\)](#)
134. [Ham radio propagation conditions](#)
135. [Ham radio propagation tutorial](#)
136. [Ham Radio Range Guide: How Far Can You Talk?](#)
137. [Ham solar](#)
138. [Helium lines](#)
139. [HF bands allocated for radio amateurs](#)
140. [HF bands allocated for radio amateurs and their characteristics](#)
141. [HF band conditions](#)
142. [HF band conditions maps](#)
143. [HF bands Info affected by solar conditions](#)
144. [HF bands Monitoring](#)
145. [HF conditions](#)
146. [HF propagation](#)
147. [HF propagation app](#)
148. [HF propagation blackout](#)
149. [HF propagation calculator](#)
150. [HF propagation chart](#)
151. [HF propagation conditions](#)
152. [HF propagation conditions at a glance](#)
153. [HF propagation experts](#)
154. [HF propagation explained](#)
155. [HF propagation explanation](#)
156. [HF propagation fadeouts](#)
157. [HF propagation for beginners and advanced radio hams](#)
158. [HF propagation for radio amateurs](#)
159. [HF propagation for radio hams](#)
160. [HF propagation forecast for radio amateurs](#)

161. [HF propagation forecast tools](#)
162. [HF Propagation Forecasting](#)
163. [HF propagation: fundamentals](#)
164. [HF propagation live map](#)
165. [HF propagation map](#)
166. [HF propagation models](#)
167. [HF propagation modes](#)
168. [HF propagation online map](#)
169. [HF propagation overview](#)
170. [HF propagation now \(real-time\)](#)
171. [HF propagation predictions](#)
172. [HF propagation prediction online](#)
173. [HF propagation prediction software](#)
174. [HF propagation preface](#)
175. [HF propagation reporter](#)
176. [HF propagation Status](#)
177. [HF propagation — The basics](#)
178. [HF Propagation Today](#)
179. [HF Propagation Tools](#)
180. [HF radio propagation indices and scales](#)
181. [HF radio propagation prediction](#)
182. [HF radio propagation tutorial for radio amateurs](#)
183. [HF propagation widget](#)
184. [HF radio conditions](#)
185. [HF radio propagation](#)
186. [HF signals propagation](#)
187. HF skywave window: LUF—MUF
188. [HF radio propagation forecast](#)
189. [HF radio propagation today](#)
190. [HF radio wave propagation](#)
191. [HF radio waves \(basics\)](#)
192. [HF radio wave propagation factors](#)

193. [HF radio propagation prediction](#)
 194. [HF real time propagation](#)
 195. [HF Skywave propagation for radio amateurs](#)
 196. [HF wave propagation](#)
 197. [HFTA - High frequency Terrain Assessment](#)
 198. [Helium lines](#)
 199. [High frequency](#)
 200. [Highest Possible Frequency \(HPF\)](#)
 201. [How skywave propagation works](#)
 202. [HPF—Highest Possible Frequency](#)
-

I

203. [IMF - Interplanetary Magnetic field](#)
204. [Impacts of space weather](#)
205. [Incident angle](#)
206. [Interplanetary Magnetic field \(IMF\)](#)
207. [Introduction to HF propagation](#)
208. [Ionisation](#)
209. [Ionization](#)
210. [Ionogram](#)
211. [Ionosonde](#)
212. [Ionosphere](#)
213. [Ionosphere and HF radio propagation](#)
214. [Ionosphere propagation of radio waves](#)
215. [Ionosphere regions](#)
216. [Ionosphere, radio waves](#)
217. [Ionosphere probing](#)
218. [Ionosphere regions](#)
219. [Ionospheric absorption—D-region](#)
220. [Ionospheric bubbles](#)
221. [Ionospheric clouds](#)

- 222. [Ionospheric conditions](#)
 - 223. [Ionospheric data](#)
 - 224. [Ionospheric disturbances](#)
 - 225. [Ionospheric irregularities](#)
 - 226. [Ionospheric propagation](#)
 - 227. [Ionospheric refraction](#)
 - 228. [Ionospheric regions](#)
 - 229. [Ionospheric skywave propagation](#)
 - 230. [Ionospheric storms](#)
 - 231. [IRI ionospheric model](#)
 - 232. [ITU model](#)
-

K

- 233. [K-Index](#)
 - 234. [Key concepts of HF propagation](#)
 - 235. [Key Factors Affecting HF Propagation](#)
 - 236. [Kp index](#)
 - 237. [KiwiSDR](#) map of remote public SDR receivers accessible via the Internet
-

L

- 238. [Lagrange points](#)
 - 239. [Learning about space weather](#)
 - 240. [Line-Of-Sight propagation](#) (LOS)
 - 241. [Live HF propagation map](#)
 - 242. [Long-distance communication](#)
 - 243. [Live solar events and past solar activity](#)
 - 244. [LOS](#)
 - 245. [LOS - Line of sight propagation](#)
 - 246. [LUF - Lowest Usable frequency](#)
 - 247. [LUF Chart - The Lowest Usable frequency map](#)
-

M

- 248. [Magnetic field](#)
 - 249. [Magnetosphere \(MS\)](#)
 - 250. [Maximum usable frequency \(MUF\)](#)
 - 251. [Maximum usable frequency map](#)
 - 252. [Minimum usable frequency map](#)
 - 253. [Modes of HF radio propagation](#)
 - 254. [Modes of radio propagation](#)
 - 255. [Monitor band activity of radio amateurs ↗](#)
 - 256. [Monitor HF propagation](#)
 - 257. [Monitor propagation conditions](#)
 - 258. [MS — Meteor Scatter Activity](#)
 - 259. [MUF — Maximum Usable frequency](#)
 - 260. [MUF Indicators](#)
 - 261. [MUF](#) How what, why?
 - 262. [MUF factor](#)
 - 263. [MUF map](#)
 - 264. [MUF propagation map by KC2G](#) for a 3,000 km hop
 - 265. [MUF propagation online map](#)
-

N

- 266. [Near real-time regional maps](#)
 - 267. [NET ionospheric model](#)
 - 268. [NOAA](#) National Oceanic and Atmospheric Administration
 - 269. [Noise level](#)
 - 270. [NVIS](#)
 - 271. [NVIS map](#)
 - 272. [NVIS propagation](#)
 - 273. [NVIS propagation concept](#)
 - 274. [NVIS tutorial](#)
-

O

- 275. [Observations - Terrestrial & Solar](#)

276. [On-Line HF propagation prediction](#)
 277. [Online propagation tools](#)
 278. [Online real-time propagation charts](#)
 279. [Optimal Sending frequency](#)
 280. [OWF - Optimum Working frequency](#)
-

P

281. [Past solar activity](#)
282. [PCA—polar cap absorption](#)
283. [Plasma](#)
284. [Plasma clouds](#)
285. [Polar cap absorption \(PCA\)](#)
286. [Practical applications](#)
287. [Predict conditions](#)
288. [Predict HF radio propagation](#)
289. [Predict MUF](#)
290. [Predict propagation](#)
291. [Predict skywave propagation now](#)
292. [Predict sunspot numbers](#)
293. [Prediction of sky-wave propagation conditions](#)
294. [Predicting propagation](#)
295. [Prediction vs forecast of HF band conditions](#)
296. [Propagation conditions](#)
297. [Propagation conditions monitoring](#)
298. [Propagation DX](#)
299. [Propagation factors and conditions](#)
300. [Propagation forecast](#)
301. [Propagation indicators](#) — A review of skywave propagation indicators
302. [Propagation indices](#) single: propagation index
303. [Propagation modes](#)
304. [Propagation of high frequency radio waves](#)

- 305. [Propagation of HF radio waves - Global conditions and Regional conditions](#)
 - 306. [Propagation of shortwave radio](#)
 - 307. [Proton flux](#)
 - 308. [Propagation | modes of HF radio propagation](#)
 - 309. [Propagation forecast](#)
 - 310. [propagation of high frequency radio waves | the rebirth of hf](#)
 - 311. [Propagation of radio waves explained](#)
 - 312. [Propagation prediction](#)
 - 313. [Propagation prediction programs and forecasts](#)
 - 314. [Propagation report](#)
 - 315. [Proton flux \(Pf\)](#)
 - 316. [PSKR - PSK reporter](#)
-

Q

- 317. [Quick guide to HF propagation using solar indices](#)
-

R

- 318. [Radio Amateur](#)
- 319. [Radio amateurs HF bands characteristics](#)
- 320. [Radio blackouts / fadeouts; R₁₋₅ scale](#)
- 321. [Radio frequency](#)
- 322. [Radio propagation](#)
- 323. [Radio propagation beacon](#)
- 324. [Radio propagation conditions today](#)
- 325. [Radio propagation forecast](#)
- 326. [Radio propagation free software](#)
- 327. [Radio propagation online tools](#)
- 328. [Radio propagation properties](#)
- 329. [Radio propagation software](#)
- 330. [Radio propagation tool](#)

331. [Radio propagation tutorial](#)
332. [Radio wave propagation in the ionosphere](#)
333. [Radio spectrum](#)
334. [Radio waves](#)
335. [Radio waves propagation](#)
336. [Radio wave propagation model](#)
337. [Radio-wave HF propagation models](#)
338. [Radio-wave propagation modes](#)
339. [Ray tracing \(Optics\)](#)
340. [RBN - Reverse beacon Network](#)
341. [Real-time activity of radio amateurs](#)
342. [Real-time band conditions](#)
343. [Real-time DX data](#)
344. [Real-time DX propagation conditions](#)
345. [Real-time HF band conditions](#)
346. [Real-Time hf propagation charts](#)
347. [Real-time HF propagation prediction](#)
348. [Ham radio HF propagation real-time](#)
349. [Real-time HF propagation reports](#)
350. [Real-time ham band activity using the internet](#)
351. [Real-time MUF 3000 km propagation map](#)
352. [Real-Time MUF and solar indices](#)
353. [Real-time MUF map](#)
354. [Real-time propagation and band conditions](#)
355. [Real-time radio propagation reports](#)
356. [Real-time reports of HF propagation](#)
357. [Real-time reports of radio propagation](#)
358. [Real-time reports of space weather](#)
359. [Real-time space weather reports](#)
360. [Real-time watching of worldwide hams' activity](#)
361. [Real-time watching of worldwide ham activity](#)
362. [Recent geophysical & solar observations](#)

- 363. [Recent MUF distribution](#)
 - 364. [Recent observations - Geo \(Terrestrial\) & Solar](#)
 - 365. [Recent sunspot number](#)
 - 366. [Regional conditions](#)
 - 367. [Regional HF conditions](#)
 - 368. [Regional MUF and solar indices](#)
 - 369. [Regional vs Global HF propagation conditions](#)
 - 370. [Regional vs Global propagation conditions](#)
 - 371. [Regions of ionization](#)
 - 372. [Region vs. Layer: Earth's Atmosphere and Ionosphere](#)
 - 373. [Remote sensing of the ionosphere](#)
 - 374. [Reports of space weather](#)
 - 375. [RSG scales of Space Weather](#)
-

S

- 376. [Satellites for space weather observations](#)
- 377. [SDO - Solar Dynamic Observatory \(NASA\)](#)
- 378. [SDR - Software Defined radio](#)
- 379. [SDR spectrum](#)
- 380. [Seasonal variations in HF radio propagation](#)
- 381. [Seasons affect HF radio propagation](#)
- 382. [SEP](#)
- 383. [SFI - Solar flux Index](#)
- 384. [Shortwave propagation](#)
- 385. [Shortwave propagation conditions](#)
- 386. [Shortwave propagation forecast](#)
- 387. [Shortwave radio propagation](#)
- 388. [SID - Sudden ionospheric disturbances](#)
- 389. [Significant frequencies relevant to skywaves](#)
- 390. [Skip distance](#)
- 391. [Skip zone](#)
- 392. [Skywave](#)

393. [Skywave concept](#)
394. [Skywave HF radio propagation](#)
395. [Skywave propagation](#)
396. [Skywave propagation concept](#)
397. [Skywave propagation conditions](#)
398. [Skywave propagation for radio amateurs](#)
399. [Skywave propagation indicators for radio amateurs](#)
400. [Skywave propagation indicators for radio hams](#)
401. [Skywave propagation forecast](#)
402. [Skywave propagation tutorial](#)
403. [Software for HF propagation prediction](#)
404. [Solar activity](#)
405. [Solar Activity & Ham radio propagation](#)
406. [Solar activity affects HF propagation](#)
407. [Solar activity affects skywaves](#)
408. [Solar activity for radio amateurs](#)
409. [Solar conditions affect HF radio propagation](#)
410. [Solar conditions and ham radio propagation](#)
411. [Solar conditions and HF radio propagation](#)
412. [Solar cycle](#)
413. [Solar Cycle Ham Radio](#)
414. [Solar data](#)
415. [Solar electromagnetic spectrum](#)
416. [Solar events](#)
417. [Solar energetic Particle \(SEP\)](#)
418. [Solar events](#)
419. [Solar flares](#)
420. [Solar flux index \(SFI\)](#)
421. [Solar flux today](#)
422. [Solar Ham](#)
423. [Solar indices](#)
424. [Solar observations](#)

425. [Solar maximum](#)
426. [Solar minimum](#)
427. [Solar Particle event \(SPE\)](#)
428. [Solar phenomena](#)
429. [Solar propagation](#)
430. [Solar radiation](#)
431. [Solar radiation storms \(flares\); S₁₋₅ scale](#)
432. [Solar radio flux](#)
433. [Solar spectra](#)
434. [Solar storms](#)
435. [Solar Storms Ham Radio](#)
436. [Solar synoptic map](#)
437. [Solar terminator](#) (i.e. Greyline)
438. [Solar wind; The impact of the solar wind on HF radio transmission.](#)
439. [Solar wind reports](#)
440. [Solar X-Ray](#) (Phenomenon)
441. [Solar X-ray Imaging system \(SXI\)](#)
442. [Solar X-Ray flux levels](#)
443. [Space wave](#)
444. [Space Weather](#)
445. [Space weather data and plots](#)
446. [Space weather definitions and explanations](#)
447. [Space weather events](#)
448. [Space weather forecasts](#)
449. [Space weather ground-based observatories](#)
450. [Space weather observations](#)
451. [Space weather observations from satellites](#)
452. [Space weather prediction](#)
453. [Space weather prediction center](#) (NOAA)
454. [Space Weather predictions for radio amateurs](#)
455. [Space weather reports](#)

456. [Space weather scales](#)
 457. [SPE](#)
 458. [Spread F](#)
 459. [Sudden ionospheric disturbances](#)
 460. [Sunspots](#)
 461. [SuperDARN](#)
 462. [SWPC = Space Weather prediction center](#) (NOAA)
 463. [Sporadic e](#)
 464. [Stratosphere](#)
 465. [Sudden ionospheric disturbances](#) (SID)
 466. [Sunlight](#)
 467. [Sunspots](#)
 468. [Sunspot cycle](#)
 469. [Sunspot number \(SSN\)](#)
 470. [Sunspots recent days](#)
 471. [Sunspots today](#)
 472. [Surface wave propagation](#)
 473. [Space Weather prediction center](#) (NOAA)
-

T

474. [T Index map](#)
475. [TEC - Total electron content](#)
476. [Terminator line](#)
477. [Terrestrial geomagnetic indices](#)
478. [Terrestrial observations](#)
479. [The Holy Cluster](#)
480. [The Rebirth of Shortwave radio communication](#)
481. [The recent HF propagation conditions](#)
482. [The recent propagation conditions](#)
483. [The recent skywave propagation conditions](#)
484. [The Super Dual Auroral Radar Network](#) (SuperDARN)
485. [Thermosphere](#)

- 486. [TID \(Traveling ionospheric disturbances\)](#)
 - 487. [Tools for analyzing and forecasting HF propagation](#)
 - 488. [Total electron content](#) — TEC
 - 489. [Transmission Angle](#)
 - 490. [Traveling ionospheric disturbances](#) (TID)
 - 491. [Troposphere](#)
 - 492. [Twilight zone](#)
-

U

- 493. [Ultraviolet](#)
 - 494. [Understanding Ham bands conditions banner](#)
 - 495. [Understanding Ham radio propagation](#)
 - 496. [Understanding HF propagation](#)
 - 497. [Understanding HF propagation review](#) 
 - 498. [Understanding HF propagation Numbers](#)
 - 499. [Understanding HF radio propagation](#)
 - 500. [Understanding LUF and MUF](#)
 - 501. [Understanding NVIS](#)
 - 502. [Understanding Skywave propagation](#)
 - 503. [Understanding the basics of HF band conditions](#)
 - 504. [Usable Frequency Range](#) from LUF to MUF
 - 505. [Usable HF frequencies](#)
 - 506. [Using HF beacon tracking programs](#)
-

V

- 507. [Variations in HF radio propagation](#)
 - 508. [Very High frequency](#) (VHF)
 - 509. [Very Low frequency](#) (VLF)
 - 510. [VHF propagation map](#)
 - 511. [View HF bands conditions at a glance](#)
 - 512. [Visualizing HF propagation](#)
 - 513. [VOACAP](#)
-

W

- 514. [Wave propagation](#)
 - 515. [What are radio waves?](#)
 - 516. [What is ham radio communications](#)
 - 517. [What is radio?](#)
 - 518. [What is radio propagation?](#)
 - 519. [What is space weather?](#)
 - 520. [What is the current LUF?](#)
 - 521. [What is the current minimum useable frequency?](#)
 - 522. [What is the current Maximum Frequency?](#)
 - 523. [What is the current MUF?](#)
 - 524. [Wide-band WebSDR](#)
 - 525. [Window of useful frequencies](#) from LUF to MUF
 - 526. [World-wide Space Weather agencies & services](#)
 - 527. [Worldwide LUF map](#)
 - 528. [WSPR - Weak Signal propagation reporter](#) * [WSPR Live](#) *
[WSPR Rocks](#) * [WSPRnet](#)
-

X

- 529. [X-Ray flares](#)
 - 530. [X-Ray flux levels](#) (scale)
-

[Total visits since 17 August 2022.](#)
Repeat visits counted as new after 24 hours.

4X4XM

US 102,590	HR 268	UG 35	VG 11	MC 3
GB 12,607	RS 268	GE 35	VI 11	CK 3
CA 8,001	HK 265	EC 33	AO 11	TG 3
DE 4,198	UA 227	GU 33	SY 11	CG 3
NL 3,335	BD 217	JM 33	BW 10	LC 3
AU 3,255	LU 206	TN 30	NE 10	AX 3
IL 2,590	LT 188	RE 30	BO 9	LI 3
IT 2,481	PK 187	SC 30	UZ 9	BZ 3
IN 2,014	SK 174	NI 30	SV 9	AD 3
FR 1,824	CO 174	OM 28	AM 9	NC 3
PL 1,584	TW 164	MM 27	CI 9	DJ 2
ES 1,542	EG 134	JO 27	PF 9	FM 2
SE 1,439	AE 130	KW 26	SR 8	GF 2
SG 1,263	IR 117	LB 26	MV 8	GL 2
BE 1,125	UY 107	NA 26	SO 7	TJ 2
JP 1,076	NG 105	BH 25	ZW 7	NF 2
BR 1,008	EE 104	PY 25	GD 7	TL 2
NO 950	IQ 98	MW 24	SN 7	KM 2
CH 922	HN 98	LY 24	CW 7	MH 2
IE 910	VN 97	TZ 24	LA 7	CM 2
ZA 880	LV 96	MU 22	MN 7	GA 2
NZ 829	SA 93	KZ 21	SD 7	TO 2
FI 793	PE 80	BB 21	PG 6	AS 2
CN 774	VE 79	KH 20	GI 6	SL 2
DK 772	MT 77	AW 20	AI 6	AG 2
PH 689	LK 70	BY 20	FK 6	GY 2
ID 682	KE 69	GH 19	MQ 6	IO 1
AT 641	ET 69	AZ 18	SM 6	KG 1
RO 598	MA 60	DO 17	DM 5	GN 1
TR 588	MZ 59	JE 17	GP 5	BQ 1
GR 561	BA 58	GG 17	VU 5	ST 1
PT 539	IS 56	BN 15	BJ 5	BI 1
CZ 507	QA 55	FO 15	SS 5	YT 1
AR 504	MD 55	AL 14	XK 5	GM 1
HU 487	CR 55	ZM 14	BT 4	TD 1
MY 471	NP 53	CD 14	ME 4	MP 1
RU 465	DZ 53	BM 14	TC 4	WS 1
MX 442	CY 52	AF 14	PS 4	CF 1
CL 381	TT 52	GT 13	BF 4	PW 1
SI 330	CU 46	VC 12	MO 4	ER 1
TH 305	MK 42	YE 12	CV 4	
BG 304	KY 41	RW 12	MG 3	
PR 303	PA 40	FJ 12	SZ 3	
KR 297	IM 39	BS 11	LR 3	

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The number of visits (graphs displayed below) peaked during HF propagation disruptions.



